



WFIRST Coronagraph Pictures for STMD

<u>Pictures collected by:</u> Nick Siegler and Bobby Effinger

<u>Pictures sourced from existing CGI presentations by:</u>

Feng Zhao
A J Eldorado Riggs
Bala Balasubramanian
John Krist
Eric Cady
Ilya Poberezhskiy
Rick Demers

August 17, 2017



Outline



- 1. Photos of coronagraph components such as masks, deformable mirrors, etc;
- Photos of the TRL-5 demonstration hardware in the JPL High-Contrast Imaging Testbed;
- 3. Photos of the labs where the masks are manufactured;
- 4. Photos of the science and engineering teams who are working on the coronagraph;
- 5. Simulated science images from the WFIRST coronagraph;
- 6. Plots to illustrate WFIRST coronagraph capabilities vs. the current state of the art;
- 7. Artists' concepts of the types of planets that the WFIRST coronagraph is designed to characterize

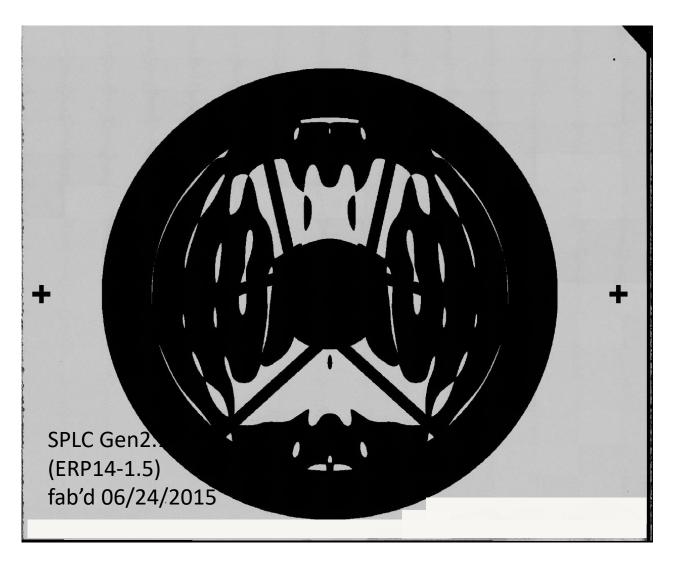




1. PHOTOS OF CORONAGRAPH COMPONENTS







Caption: WFIRST Coronagraph Shaped Pupil Planet Imaging and Spectroscopy Mask materials are Black silicon and Silver/Aluminum?

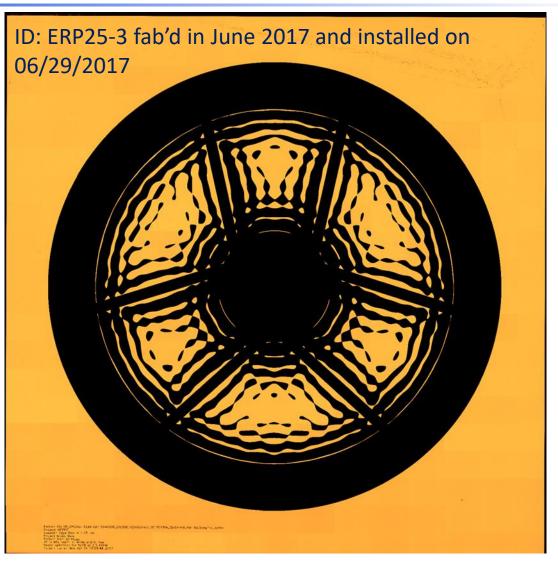
Mask is fabricated at JPL's Microdevices Laboratory





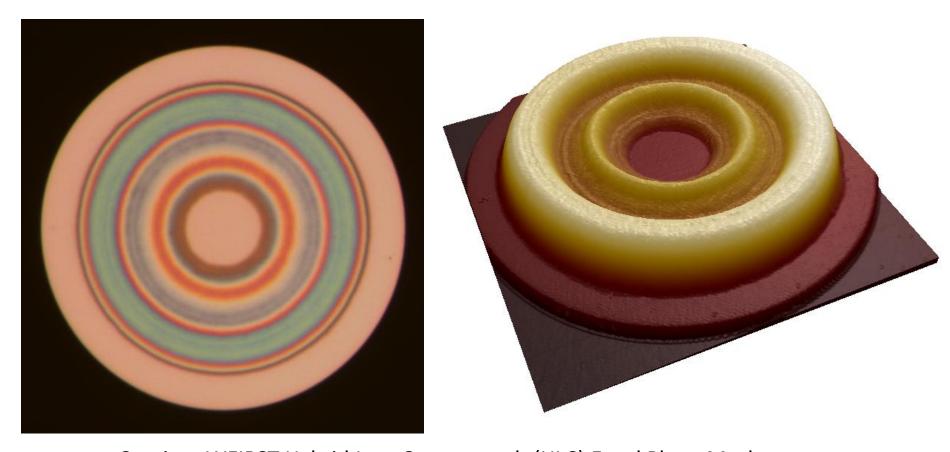


Caption: WFIRST Coronagraph Shaped Pupil Disk Science Mask in it's Optical Mount



Caption: WFIRST Coronagraph Shaped Pupil Disk Science Mask





Caption: WFIRST Hybrid Lyot Coronagraph (HLC) Focal Plane Mask

Diameter of the mask is approximately 85 microns!

Masks are nanofabricated at JPL's Microdevices Laboratory





Deformable Mirrors

- 48x48 actuators
- 48 mm diameter active area
- PMN = Lead Magnesium Niobate Pb(Mg_{1/3}Nb_{2/3})O₃
- Made by Northrup Grumman Xinetics in collaboration with JPL







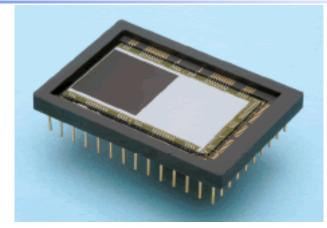
1k x 1k Electron Multiplying CCD (EMCCD)

CCD201-20 by Teledyne e2v

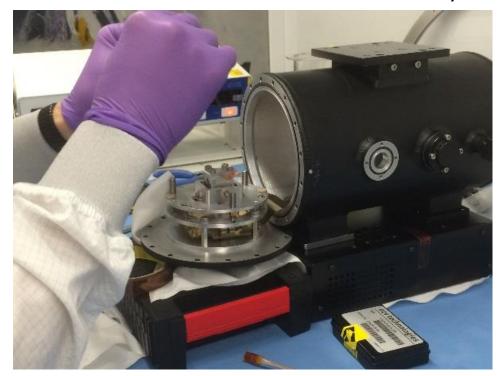
Detectors have passed radiation testing

Will be the first ultra-low noise detectors flown

in space



Caption: Performance testing of EMCCD
Detector at JPL's Low Flux Detector Laboratory





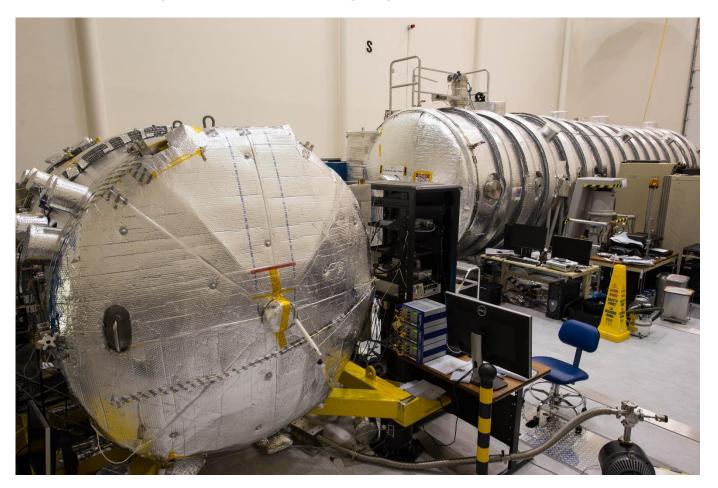


2. PHOTOS OF THE TRL-5 DEMONSTRATION HARDWARE IN HCIT





ExEP High Contrast Imaging Testbed (HCIT) Facility



Caption: HCIT-1 (left) is the vacuum chamber at JPL testing the WFIRST coronagraph; HCIT-2 (right) is testing the WFIRST coronagraph spectrometer.

2. Photos of the TRL-5 demonstration hardware in the JPL High-Contrast Imaging Testbed



ExEP High Contrast Imaging Testbed (HCIT) Facility

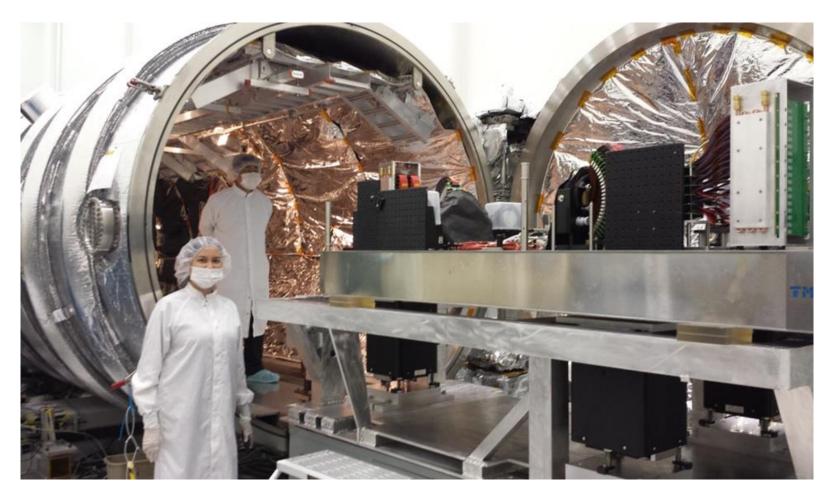


Caption: The WFIRST coronagraph inside a vacuum chamber at JPL.





ExEP High Contrast Imaging Testbed (HCIT) Facility

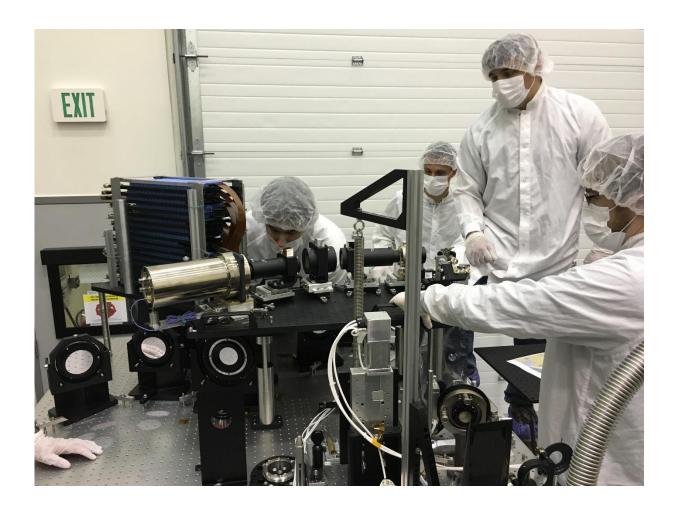


Caption: The WFIRST Hybrid Lyot coronagraph being placed inside its vacuum chamber at JPL.





ExEP High Contrast Imaging Testbed (HCIT) Facility



Caption: The Goddard Space Flight Center spectrometer (PISCES) being prepared to be inserted into its vacuum chamber at JPL.



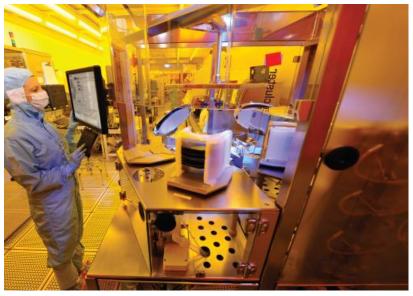


3. PHOTOS OF THE LABS WHERE THE MASKS ARE MANUFACTURED



Click for video





Caption: Mask fabrication with e-beam facility at JPL's Microdevices Laboratory





4. PHOTOS OF THE SCIENCE AND ENGINEERING TEAMS WHO ARE WORKING ON THE CORONAGRAPH



4. Photos of the science and engineering teams who are working on the coronagraph



<u>Coronagraph Adjutant Scientist</u> (CAS): Jeremy Kasdin, Princeton University <u>Science Investigations Teams</u> (SIT):

SIT #1 PI: Bruce Macintosh, Stanford University

SIT #2 PI: Maggie Turnbull, SETI Institute





4. Photos of the science and engineering teams who are working on the coronagraph





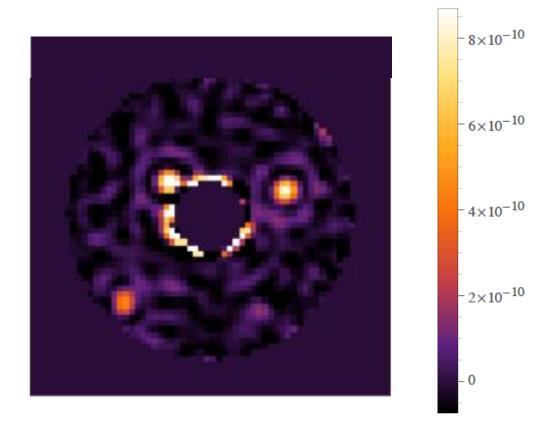




5. SIMULATED SCIENCE IMAGES FROM THE WFIRST CORONAGRAPH

5. Simulated science images from the WFIRST coronagraph

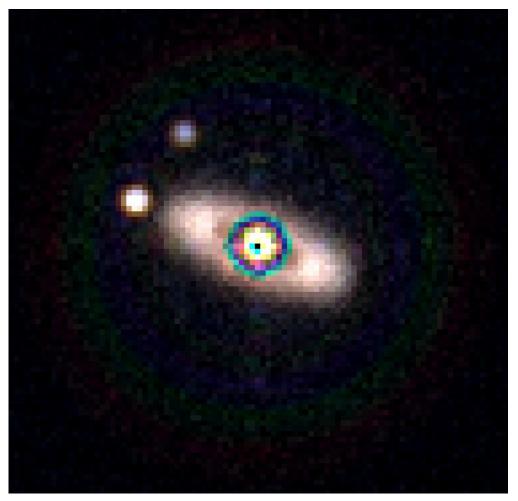




Caption: Simulated image of science target 47 UMa using the WFIRST Hybrid Lyot Coronagraph. A post-processing algorithm, called KLIP, was used to pull the planet signal from the noisy background light.



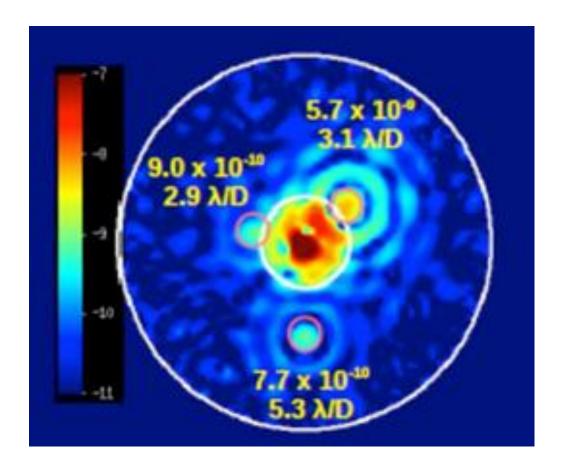




Caption: Simulated image of 1 zodi circumstellar disk along with two gas giants exoplanets (Jupiter and Staurn analogs). Picture taken from Exo-C STDT final report.







Caption: Simulated image of three gas giants exoplanets using the WFIRST Hybrid Lyot Coronagraph.





6. PLOTS TO ILLUSTRATE WFIRST CORONAGRAPH CAPABILITIES VS. THE CURRENT STATE OF THE ART





Caption and Notes for the next slide

Notes:

This is the money slide. This chart, if understood, is the reason the WFIRST coronagraph is so important to NASA's search for life effort. It is a key technology stepping stone from the SOA (HST and ground telescopes) to the capability we need to directly image exo-Earths and look for signs of life. WFIRST coronagraph advances the SOA from HST by 4 orders of magnitude and the ground by at least 2 orders. That leaves us 1-2 orders away from the requirements for the next NASA exoplanet mission.

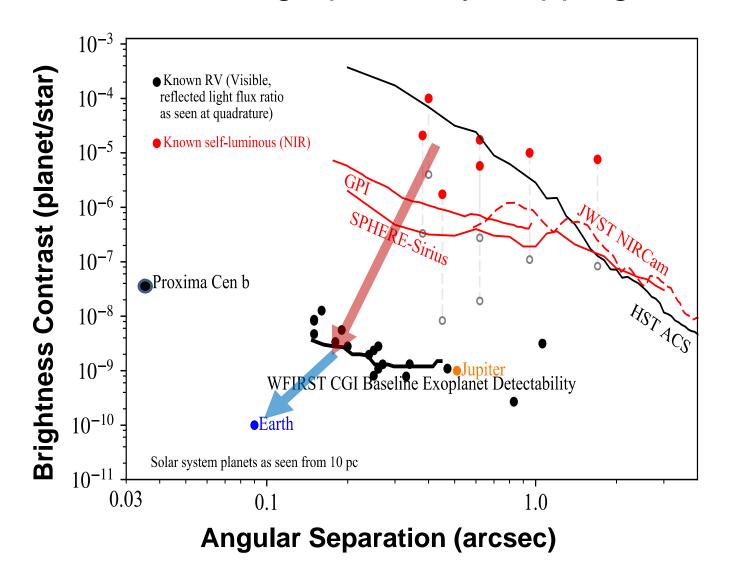
Caption:

➤ Brightness contrast ratio (ratio of planet brightness to host star brightness) versus apparent angular separation. The filled orange circles indicate the direct imaging of young, self-illuminous planets imaged in the near-infrared by ground-based telescopes. Contrast for the Earth point is for analogous exo-Earth placed 10 pc away. The solid black dots are contrast estimates of known radial velocity planets, including Proximus Centauri b, the closest exoplanet to the Earth. The orange curves show measured performance of ground-based coronagraphs: the GPI curve shows typical performance, while the SPHERE curve shows the best achieved performance to-date on the reference star Sirius. Achieved performance with HST/ACS coronagraphic masks, and the predicted performance of JWST/NIRCam masks are also shown. The current best estimate for WFIRST coronagraph (solid back) is at 565 nm and includes a factor of 10 contrast improvement from data post-processing. The WFIRST coronagraph contrast requirement is 3x10⁻⁹.





WFIRST Coronagraph, A key stepping stone







7. ARTISTS' CONCEPTS OF THE TYPES OF PLANETS THAT THE CGI IS DESIGNED TO CHARACTERIZE 26



7. Artists' concepts of the types of planets that the WFIRST coronagraph is designed to characterize









